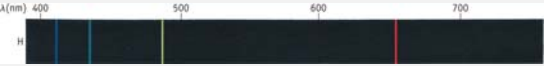


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[TP] What electron cloud energies account for the 486 nm line in the gas discharge spectrum (Balmer series) of H atoms?



17% 1. Only the $n = 3$ cloud energy
 17% 2. Only the $n = 4$ cloud energy
 17% 3. Only the $n = 5$ cloud energy
 17% 4. The $n = 2$ and $n = 4$ cloud energies
 17% 5. The $n = 2$ and $n = 5$ cloud energies
 17% 6. None of these

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Response Counter 10 1

Lecture 31 CH101 A1 (MWF 9 am)
 Monday, November 28, 2016

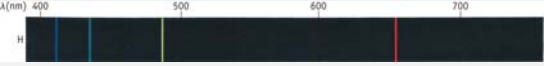
- Complete: H atom photon energies
- He^+ , Li^{2+} , etc., photon energies
- Photoionization (photoelectric effect)

Next lecture: Hydrogen atom electron clouds <http://goo.gl/XPkcxv>;
 Review of electron clouds; Orbital (yikes!) approximation

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What about clouds made from a single n ?

It turns out that clouds made from a single electron wave do not move.
 That they do not move is why the electrons in an atom **do not collapse** into the nucleus, and so **why atoms exist!**

It is only clouds resulting for the mixture of different electron waves that move, and it is interaction with light that creates such mixtures.

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6

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Light-matter resonance questions

Question: We have said electrons in an atom are a **non-moving cloud**. Then, how can there be “jiggling” at the light frequency?

Answer: Clouds for electron energy states indeed **do not** “jiggle”.

Rather, when light interacts with matter, it produces a **mixture of electrons states** of energy E_i and E_f .

It is these mixtures whose clouds “jiggle” at the frequencies

$$\nu_{\text{light}} = \nu_{\text{cloud}} = |E_f - E_i| / h$$



7

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H atom photon energies

Balmer’s formula always involves $E_2 = -Ry / 2^2$.

What about extending the formula in terms of $E_1 = -Ry / 1^2$ as

$$\Delta E_{\text{light}} = E_{\text{photon}} = h c / \lambda = Ry \left(\frac{1}{1^2} - \frac{1}{n^2} \right), n = 2, 3, 4, 5, \dots, \infty$$

Where in the spectrum would lines given by this formula occur?

UV

It turns out the lines are there!



8

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H atom photon energies

Balmer’s formula always involves $E_2 = -Ry / 2^2$.

What about extending the formula in terms of $E_3 = -Ry / 3^2$ as

$$\Delta E_{\text{light}} = E_{\text{photon}} = h c / \lambda = Ry \left(\frac{1}{3^2} - \frac{1}{n^2} \right), n = 4, 5, 6, 7, \dots, \infty ?$$

Where in the spectrum would lines given by this formula occur?

IR

It turns out the lines are there!



9

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H, He⁺, Li²⁺, etc., photon energies

The H atom expressions for E_{photon} also work for all **one-electron ions**, when multiplied by the **square of the nuclear charge** of the ion,

$$E_{\text{photon}} = h c / \lambda = Ry Z^2 \left(\frac{1}{1^2} - \frac{1}{n^2} \right), n = 2, 3, 4, \dots$$

$$E_{\text{photon}} = h c / \lambda = Ry Z^2 \left(\frac{1}{2^2} - \frac{1}{n^2} \right), n = 3, 4, 5, \dots$$

$$E_{\text{photon}} = h c / \lambda = Ry Z^2 \left(\frac{1}{3^2} - \frac{1}{n^2} \right), n = 4, 5, 6, \dots$$

etc.



16

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Ionization

What happens if an H atom absorbs a photon of energy **greater than**

$$E_{\text{photon}} = E_{\text{cloud},\infty} - E_{\text{cloud},1} = (0 + \text{Ry}) = \text{Ry}?$$



26

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Ionization

The electron is **ejected** from the atom, leaving behind H^+ .

This is called **ionization**.

$E_{\text{cloud},\infty} - E_{\text{cloud},1} = \text{Ry}$ is called the **ionization energy, IE** .



28

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Ionization

Since $\Delta E_{\text{atom}} = -\Delta E_{\text{light}}$ and energy ΔE_{light} has been **lost by the light**, the energy **in excess of the ionization energy** is carried away as **electron kinetic energy, KE** ,

$$\Delta E_{\text{atom}} = IE + KE$$



29

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Ionization

Sketch the energy diagram for ionization of an electron from Li^{2+} in its $n = 6$ level by light of wavelength 310 nm.

Use your sketch to get the **algebraic expression** for the **kinetic energy** of the electron, using h , c , Ry , and **310 nm** in your expression.



30

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Ionization

Answer: The ionization energy of Li^{2+} ($n = 6$) is

$$IE = E_{\text{cloud},\infty} - E_{\text{cloud},6} = 0 + \text{Ry } 3^2/6^2 = \text{Ry}/4$$

The photon energy is $E_{\text{photon}} = hc/(310 \text{ nm})$

The kinetic energy of the electron is

$$\begin{aligned} KE &= E_{\text{photon}} - IE \\ &= hc [1/(310 \text{ nm}) - \text{Ry}/4] \\ &= 9.6 \times 10^{-20} \text{ J} = 0.60 \text{ eV} \end{aligned}$$



31